

# Detection of circular polarization in light scattered from photosynthetic microbes

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**The identification of a universal biosignature that could be sensed remotely is critical to the prospects for success in the search for life elsewhere in the universe. A candidate universal biosignature is homochirality, which is likely to be a generic property of all biochemical life. Because of the optical activity of chiral molecules, it has been hypothesized that this unique characteristic may provide a suitable remote sensing probe using circular polarization spectroscopy. Here, we report the detection of circular polarization in light scattered by photosynthetic microbes. We show that the circular polarization appears to arise from circular dichroism of the strong electronic transitions of photosynthetic absorption bands. We conclude that circular polarization spectroscopy could provide a powerful remote sensing technique for generic life searches.**

homochirality | life detection | remote sensing

The search for life in the Universe depends on the identification of observable signatures that are unique to biological processes. If these signatures may be sensed remotely, then extensive surveys of planetary surfaces and distant objects may be undertaken without the need (initially) for costly landing spacecraft. A candidate universal biosignature that may lend itself to remote sensing application is homochirality, which, because of the optical activity of biological molecules, is potentially detectable using circular polarization spectroscopy (1–6). Organic molecules typically exist in 2 mirror-image forms and are said to be “chiral”; that is, they exhibit handedness. All known living organisms use only left-handed or l-amino acids in proteins and right-handed or d-sugars in nucleic acids and this unique preference for just a single handedness is termed “homochirality.” Intriguingly, analysis of the Murchison meteorite has shown l-excesses of 2–9% for a number of  $\alpha$ -methyl amino acids (7), with slightly smaller excesses found in the Murray meteorite (8). Homochirality is thought to be generic to all forms of biochemical life as a necessity for self-replication (9) and hence it is likely to be a signature of nonterrestrial life. To be detectable remotely using circular polarization, homochirality must imprint itself upon the circular polarization spectrum in scattered light. Here, we report on the results of sensitive laboratory measurements of the polarization spectra of light both scattered from photosynthetic microbes and in transmission through the same cultures. For context, we also present polarization spectra of a leaf and a mineral.

There is a vast array of experimentation that may be brought to bear in the case of in situ tests for the presence of biological processes (10); however, in situ experiments can sample only a tiny fraction of a planetary surface and its immediate subsurface, and they often anticipate a degree of specificity in the biology sought. Few remote sensing methods directly probe signatures of biological life. Trace gases can be observed that could have a biological origin, such as recent detections of localized methane production on Mars (11). Jupiter’s moon Europa is strongly suspected to host a liquid water ocean (12), and infrared

spectroscopic features have been shown to be consistent with those of radiation tolerant microbes (13). Beyond the Solar System, methods will be needed to assess whether extrasolar planets harbor life, and remote sensing is a necessity. Attention is being given to atmospheric composition disequilibria and to biological pigmentation spectral features as biomarkers (14–16). Typical disadvantages of these methods include model dependence and the possibility that the “biosignature” could be produced by abiotic processes, leading to a false positive.

Circular polarization may provide a more direct indication of the presence of biological processes because it is directly attributable to the chirality of the organic molecules. Earlier experiments (1, 3, 4) looked at leaves and found significant circular polarization. Here, we focus on light reflected from photosynthetic bacterial cultures. Photosynthetic life must reside at the surface, use windows of atmospheric transparency and exploit regions of the spectrum where the host star shines brightly, and hence such life forms are maximally observable. The strong electronic absorption bands that are characteristic of photosynthesis are known to exhibit circular dichroism (different absorption coefficients for left- and right-circularly polarized light); hence, we may anticipate a consequent polarization signature in scattered light, although because of the complexities of the scattering process, this is not entirely obvious in advance. For example, multiple scattering tends to randomize the polarization state; or alternatively, any circular polarization produced in the incident direction might be cancelled by circular polarization produced by subsequent reflection, because this will have the opposite handedness. If the scatterers depolarize the light, then an appropriate balance between the scattering coefficient and the differential absorption coefficient is needed to achieve measurable circular polarization. In this work, we are testing whether that balance can be achieved with microorganisms. Circular polarization can also be caused by optical interaction associated with the chirality of subcellular structures, such as membranes and macromolecules, aspects that clearly relate to the presence of biology, although yielding a distinct spectral signature relative to the circular dichroism of absorption bands (17).

Photosynthetic cyanobacteria arose between 2 and 3 billion years ago (18, 19). The enormous evolutionary advantages of photosynthesis coupled to the resilience of these microbes led to planet-wide changes in the demographic abundance of terrestrial life forms and in turn, to major changes in the composition of the

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Earth's atmosphere, specifically, to the rise of oxygen. That this happened early in Earth's history suggests that in an astrobiological context the occurrence and success of early microbial photosynthesis is plausibly commonplace, long-lived and statistically dominant on randomly chosen extrasolar planets. Therefore, it is possible that we may find extrasolar planets populated with organisms not unlike, although presumably not identical to, those we study here. The evolutionary advantages of photosynthesis also suggest that if life elsewhere in the Solar System has succeeded in adapting to the harsh extremes of a surface environment, then it too will likely use the energy consequently available through photosynthesis and phototrophy.

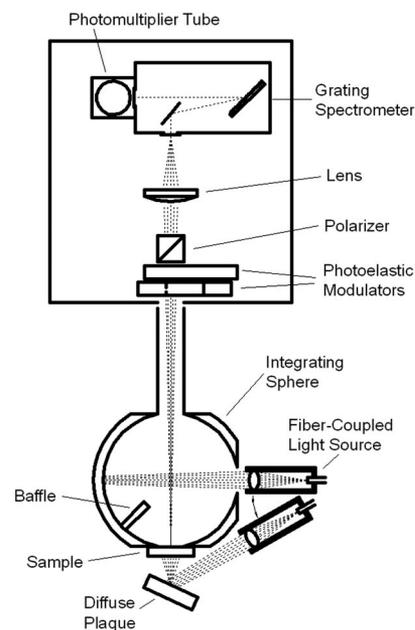
We therefore selected and cultured 2 strains of hardy chlorophyll-based photosynthetic marine cyanobacteria (20) and an  $\alpha$ -proteobacterium that uses a different phototrophic apparatus (the bacteriochlorophyll of nonsulfur purple bacteria). A variety of antenna pigments are represented. Unicellular cyanobacteria of the genus *Synechococcus* are among the most abundant members of the picophytoplankton (20–23) and contain phycobilisome accessory light harvesting pigments. *Synechococcus* WH8101 appears green because of the presence of phycocyanin (PC), whereas *Synechococcus* WH7805, rich in phycoerythrin (PE), appears pink. The purple nonsulfur bacteria are of interest because of their capacity for anaerobic, nonoxygenic photosynthesis and growth with hydrogen production. We chose a versatile  $\alpha$ -proteobacterium, *Rhodospirillum rubrum* strain ATCC 11170, possessing a well-characterized carotenoid light-harvesting apparatus, surrounding bacteriochlorophyll photo-centers (24, 25).

## Results

Liquid cultures of microbes were placed horizontally in a shallow glass Petri dish, exposed to the air, approximately at room temperature. To investigate a potential connection to circular dichroism, the experiment was configured to acquire both transmission and scattered light reflection polarization spectra, using a dedicated polarimeter (see *Materials and Methods*). The sample was illuminated from above with diffuse unpolarized light for the scattering experiments. For transmission measurements, the samples were illuminated from behind by the same white light that shone onto a diffuse white plaque beneath the sample. A specialized polarimeter viewed the sample directly down the surface normal to measure the polarization of the scattered or transmitted light returned from the sample (Fig. 1).

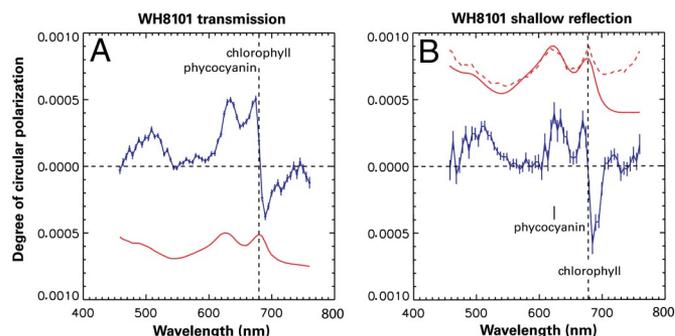
Figs. 2–4 show the polarization spectra for our samples with 1 SD error bars. Scaled absorption spectra are overlaid to reveal any relationships between absorption bands and circular polarization (the Cotton effect (26, 27)). The cyanobacteria (Figs. 2 and 3) show chlorophyll *a* absorption in the red ( $\approx 680$  nm) and the antenna pigments phycocyanin ( $\approx 620$  nm) and phycoerythrin ( $\approx 560$  nm) absorption bands as expected. In Fig. 3A, the blue chlorophyll *a* absorption band at  $\approx 430$  nm is also visible. The polarization and absorption spectra of *R. rubrum* are shown in Fig. 3B. The primary photosynthetic reaction center molecule, bacteriochlorophyll, is seen with its absorption peak at  $\approx 590$  nm, and with its 800-nm peak at the very edge of the measurement window. A mix of carotenoids causes absorption in the blue regions of the spectra. Chlorophyll *a* is absent from the *R. rubrum* spectra.

The transmission circular polarization spectra are analogous to a classical circular dichroism experiment used in protein structure and conformation analysis (28). Hence, we expected to see Cotton effect circular dichroism signatures of the strong electronic absorption features and this was indeed the case. For WH8101 (Fig. 2A), we see a broad absorption complex in the blue (carotenoids), and bands at 620 nm (phycocyanin) and 680 nm (chlorophyll *a*). Each shows a significant circular polarization signal. Furthermore, whereas the blue complex and the phycocyanin band display circular polarization of single sign, the circular polarization of the chlorophyll *a* band displays a very distinctive derivative-shaped “conservative” circular dichroism signature (29–31). The sign of the circular polarization reverses precisely at the location of the absorption maximum. This well-known effect is due to the presence of exciton-coupled chlorophyll molecular dimers where chlorophyll molecules in close proximity to one another function in pairs, effectively acting as a macromolecule.



**Fig. 1.** Illustration of the experimental configuration. In reflection mode, the light from a fiber-coupled quartz-tungsten-halogen lamp enters through an open port (B) into a 200-mm diameter integrating sphere, illuminating a spot on the sphere wall opposite. A baffle is located between the illuminated spot and the sample at port A to reduce direct illumination from that direction. The light is depolarized by numerous internal reflections and exits as unpolarized diffuse light onto the surface of the sample at port A. The polarimeter views the sample from the surface normal at port C and the sample at port A is imaged onto the entrance slit of the monochromator. In transmission mode, the horizontal port B is closed and the same light source is used to illuminate a diffuse white plaque beneath the sample and this low-polarization white light passes through the sample directly to the polarimeter.

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**Fig. 2.** Circular polarization spectra of cyanobacteria WH8101. (A) Transmission polarization spectrum of *Synechococcus* WH8101. The blue line shows the degree of circular polarization, with  $\pm 1$  SD error bars; the solid red line shows a scaled version of the absorbance spectrum. (B) Reflection polarization spectrum as in A except that the solid red line is scaled  $-\log_{10}(\text{Reflectance})$ , and the dashed red line is a scaled plot of linear polarization degree.



spectra of low amplitude and different character to biological ones (1, 2, 33). Kemp et al. (2) present circular polarization spectra of many of the Solar System bodies. A variety of effects result in polarization degree in actual abiotic situations of order  $10^{-5}$  typically. The spectral dependence is smooth and slowly varying, e.g., as atmospheric scattering gives way to surface scattering on Mars from blue to red, or as in the geometric chirality of the polar effect in gas giant planets where opposite hemispheres exhibit polarization of opposite sign from scattering. The circular polarization arising from dielectric and metallic powders was investigated empirically and theoretically (33), and no significant polarization was found from the dielectric material, whereas the metallic powders produced circular polarization from multiple scattering. This caused variations with phase angle (viewing direction relative to incident direction) and may prove to be a useful diagnostic for metal-rich objects such as certain asteroids. The lack of pronounced spectral features and low polarization amplitude is typical of these studies, although a more extensive study of abiotic scattering coupled to empirical field polarization observations in natural environments is required to be fully confident that false positives are rare. Nevertheless, these initial indications are encouraging.

On Earth, it is reasonable to expect that densely vegetated regions will produce a significant polarization signature. More challenging is the oligotrophic ocean, from which cyanobacterial sample WH7805 was derived. The oceans present a wide range of chlorophyll content, of order  $0.01\text{--}50\text{ mg}\cdot\text{m}^{-3}$ . Ocean scattering and reflectance optics are dominated by chlorophyll when its concentration exceeds  $\approx 0.14\text{ mg}\cdot\text{m}^{-3}$  (34, 35). The chlorophyll concentration is dominated by the phytoplankton biomass, and this optical transition corresponds to a cyanobacteria density of  $\approx 3 \times 10^{10}$  cells per cubic meter (36). The corresponding euphotic depth is 100 m and less for higher concentrations (35). Our heuristic interpretation of the observation that the scattered and transmitted polarization levels are comparable (Figs. 2 and 3) is that the polarization is produced within the layer above optical depth unity, or  $\approx 0.2 \times$  the euphotic depth. Hence, for oligotrophic oceanic regions whose spectral properties are dominated by the scattering and absorption from phytoplankton chlorophyll, we anticipate an implicit polarization level comparable to that of the laboratory measurements that were optimized so that the transmission optical depth  $\tau \approx 1$ , by dilution of a pellet containing few  $\times 10^{10}$  cells. Extraneous dilution of the oceanic polarization signal will undoubtedly occur because of the presence of other particulate scatterers (although many of those will be biological), surface roughness, atmospheric scattering, and clouds. Nevertheless, ocean color can be dominated by chlorophyll and hence phytoplankton, offering the possibility that a circular polarization degree of order  $10^{-4}$  or higher may be present. Empirical measurements are clearly needed.

Elsewhere within the Solar System the ability to achieve high spatial resolution, and hence lower dilution, at high light level indicates that planetary surface polarization surveys would be feasible. For example, a detailed survey of the surface of Mars for chiral spectropolarization signatures would be possible. Circular polarization imaging of a portion of the Mars surface at 2 wavelengths was carried out (6) with 210-km spatial resolution, using the European Southern Observatory Very Large Telescope. Although covering only a small fraction of the surface and with extremely limited wavelength coverage, their null results are encouraging as a proof of observational concept and by the absence of any false positives.

Although the prevailing opinion is that the surface of Mars is too hostile for life, Landis (37) describes the characteristics that would be required for a viable microorganism on Mars and shows that there are terrestrial examples in the halobacteria (38). Also, very commonly terrestrial microbial life is found not highly dispersed, but in tight knit localized colonies, films, and mats for

protection and survival and to take advantage of localized niches of habitability, because of the presence of moisture and nutrients acceptable to the microbes (39, 40). Complete spectral coverage and high spatial resolution would be required to probe for localized surface microbial communities. Additionally, a potentially long racemization timescale at the surface of Mars (41) offers the possibility of seeking fossil evidence of the remnants of long-extinct biological activity, using indicators of chirality. Other potential Solar System targets include Europa and Titan, and more primitive bodies such as comets and asteroids (42).

Eventually, with the advent of ground-based “Extremely Large Telescopes” and future dedicated space missions such as NASA’s Terrestrial Planet Finder (TPF) and beyond, it will become possible to detect and characterize Earth-like extrasolar planets. The presence of biological pigments and the distinctive red-edge of chlorophyll have been proposed as potentially useful biosignatures (15, 16). Circular polarization may provide an important complement to the use of chlorophyll’s red edge and in discerning whether an apparent pigmentation spectral feature has a biological origin, provided sufficient photons can be accumulated. It is unlikely that a first-generation space-based TPF mission would be able to collect enough photons for circular polarimetry, however, some of the larger ground-based telescopes under consideration, such as the European Extremely Large Telescope 42-m concept, would have sufficient light gathering capacity to carry out circular spectropolarimetry on Earth-like planets around some Solar neighborhood stars. (Although there will be serious issues of whether the images of these planets can be adequately separated from those of their host stars.) Beyond oxygenic photosynthesis used by cyanobacteria and implied in the use of chlorophyll’s red edge as a biomarker, it would be possible to detect enrichment of a planet by anoxygenic photosynthetic organisms. There is evidence that, on the early Earth, anoxygenic photosynthetic bacteria including those that use hydrogen, hydrogen sulfide, and reduced iron as electron donors preceded oxygenic photosynthesis and dominated ocean photosynthesis for more than a billion years (43), before the rise of oxygen in the Earth’s atmosphere.

It is also plausible to consider an analogue to the circular polarization transmission measurements in which the light from the host star offers a probe through the atmosphere of a transiting planet. Hence, the presence of biological molecules in the planet’s atmosphere could be revealed. This may become feasible when large amplitude transits are found from planets orbiting small dwarf stars, although circular polarization from stellar magnetic fields will add to practical implementation issues for these stars.

Hence, there are situations in which the distinctive character of circular polarization spectra—the Cotton effect correlation of circular polarization with absorption bands and distinctive sign change—could provide a very powerful indicator of the presence of optically active chiral molecules, and we conclude that under the right circumstances, circular polarization spectroscopy could be a very important tool in the search for extraterrestrial biological processes.

## Materials and Methods

**Polarimeter.** The polarimeter (Hinds Instruments, Series II/F542–47) is a dual photoelastic modulator (PEM) precision optical polarimeter optimized for measurement of circular polarization in the presence of significant linear polarization. If  $(I, Q, U, V)$  are the usual Stokes parameters, the design goal of measuring degree of circular polarization  $p_c \equiv V/I = 10^{-4}$  in the presence of linear polarization degree  $p_l \equiv \sqrt{(Q^2 + U^2)}/I = 0.03$  is achieved. Light encounters 2 PEMs oriented  $45^\circ$  to one another, modulated at resonance frequencies 42 kHz and 47 kHz, followed by a Glan prism analyzer, axis  $22.5^\circ$  to the modulators, a field lens, monochromator and photomultiplier detector. The entire system is controlled by a dedicated desktop computer. Linear Stokes parameters,  $Q$  and  $U$ , are measured by the  $2f$  modulation frequency of the



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